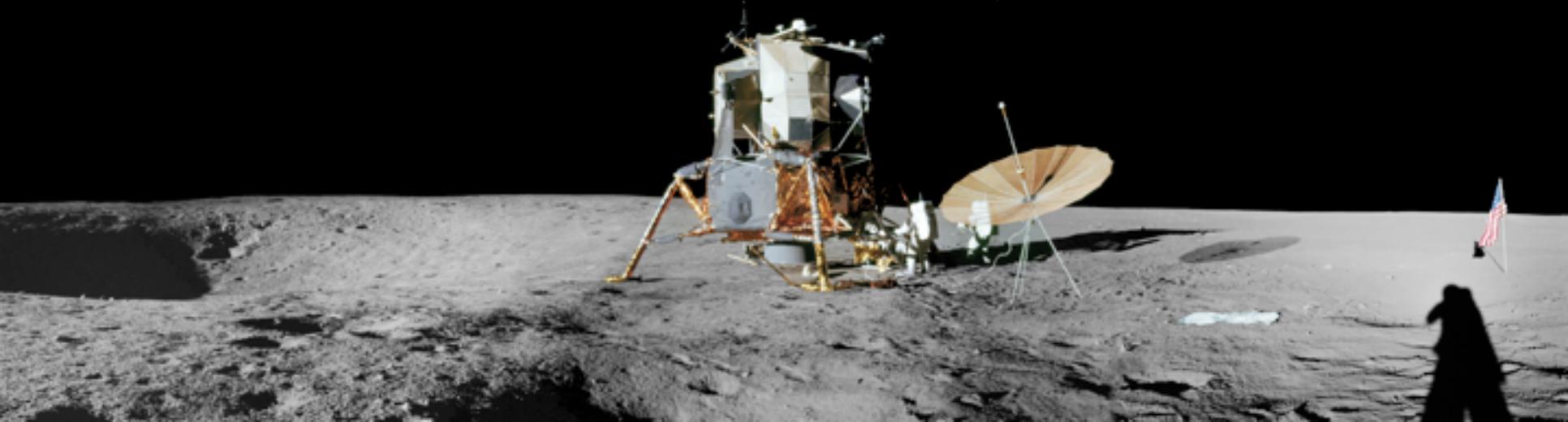


Comparison of Parental Basaltic Liquids at the Apollo 12 Site

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Purpose of Study

- Mare basalts can be used as probes into the Lunar interior
- Can gain insight into the composition of the interior
- Understand post-magma generation processes (assimilation, fractional crystallization)

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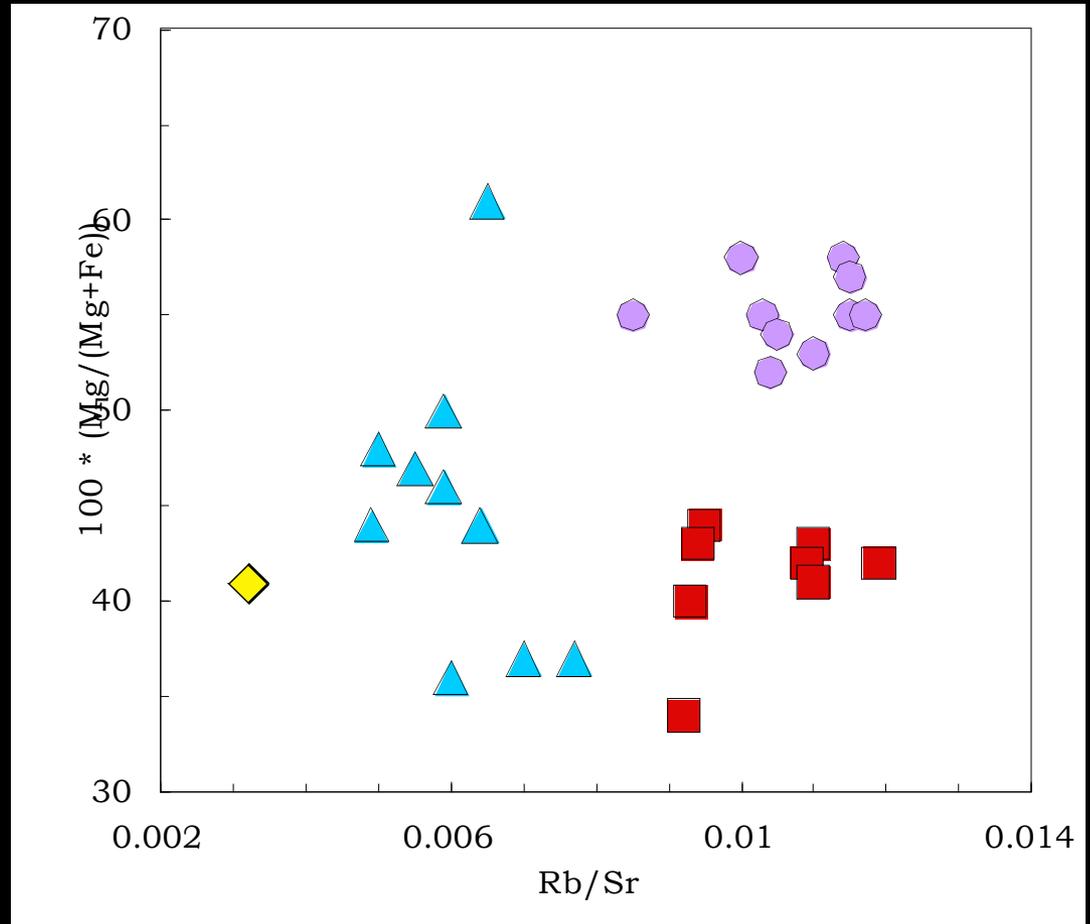
- Efficiency of Lunar Magma Ocean crystallization (overturn, fractional crystallization)
- Lunar Magma Ocean crystallization processes may apply to many planetary bodies

Background

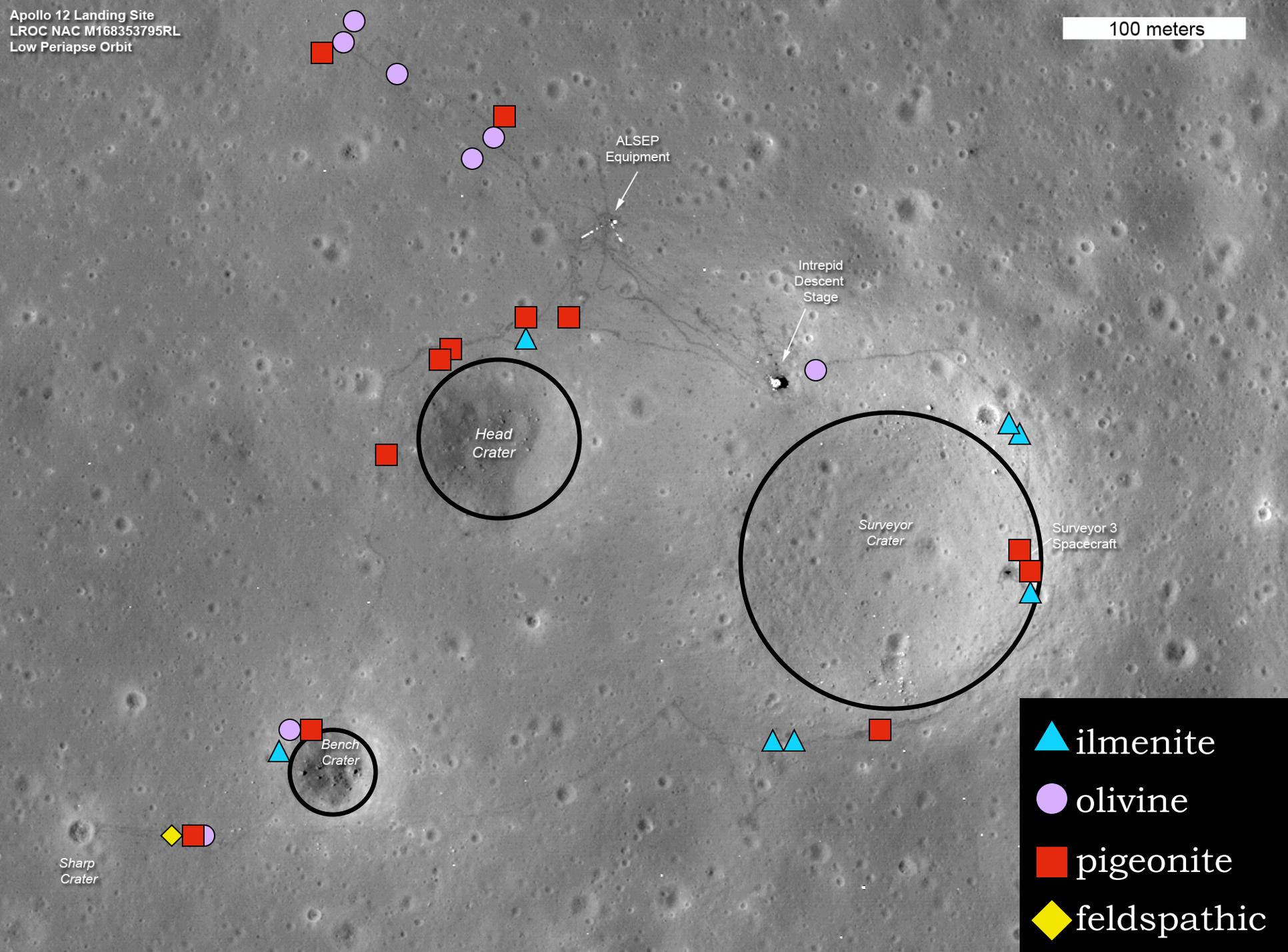
- Previous studies have determined that Apollo 12 basalts must have had multiple source regions (e.g. Dungan and Brown, 1977; Neal et al., 1994; Nyquist et al., 1979).
- Apollo 12 basalt sources isolated early in lunar history (Nyquist et al., 1979).
- The ultimate aim of this study is to determine the trace element.
- characteristics of these reservoirs and any post magma generation processes - start by defining parental melts.

Apollo 12 Basalt Types

- Based on whole rock geochemistry basalts fall into 4 groups:



Adapted from Neal et al., 1994



ALSEP
Equipment

Intrepid
Descent
Stage

Head
Crater

Surveyor
Crater

Surveyor 3
Spacecraft

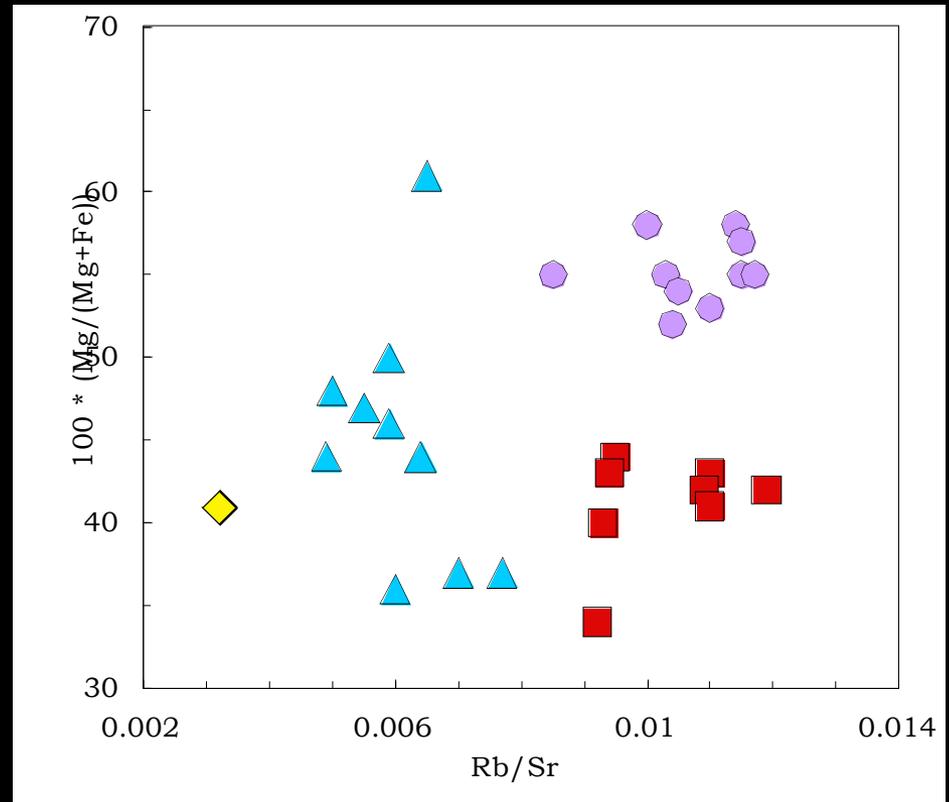
Bench
Crater

Sharp
Crater

- ▲ ilmenite
- olivine
- pigeonite
- ◆ feldspathic

Apollo 12 Basalt Types

- Our study utilizes pyroxene phenocrysts, from three of these suites (Olivine Basalts = future work):
Basalts = future work):

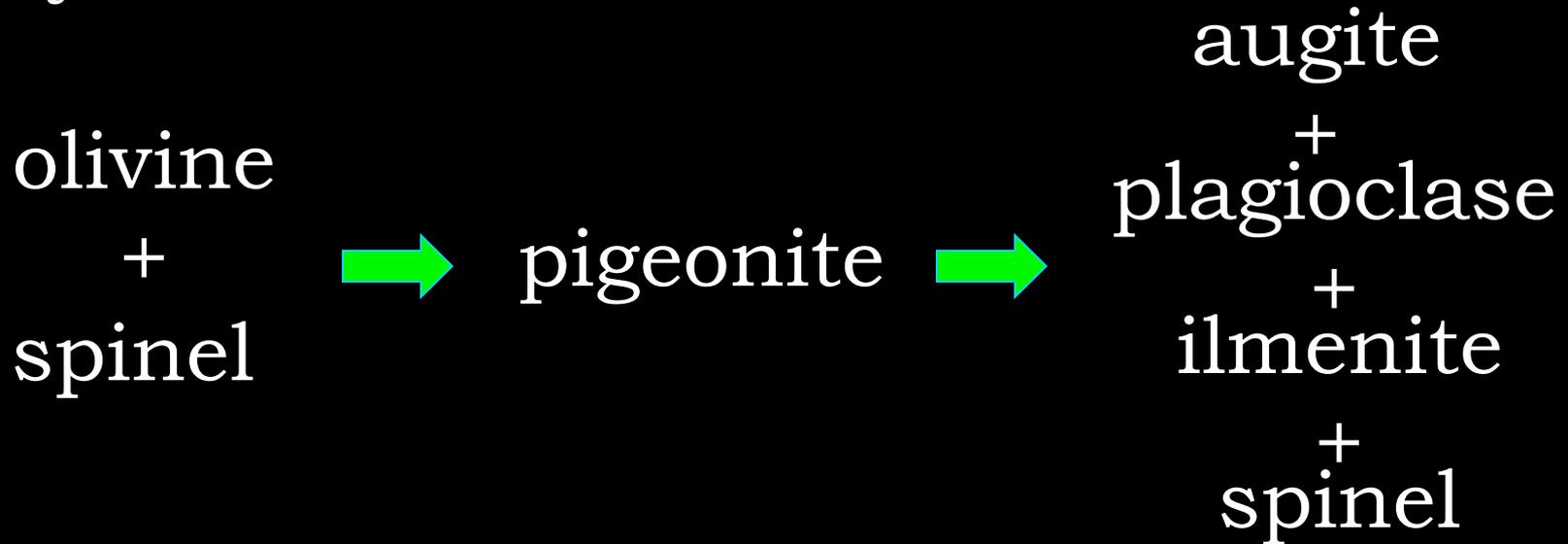


Adapted from Neal et al., 1994

Papike et al., 1979

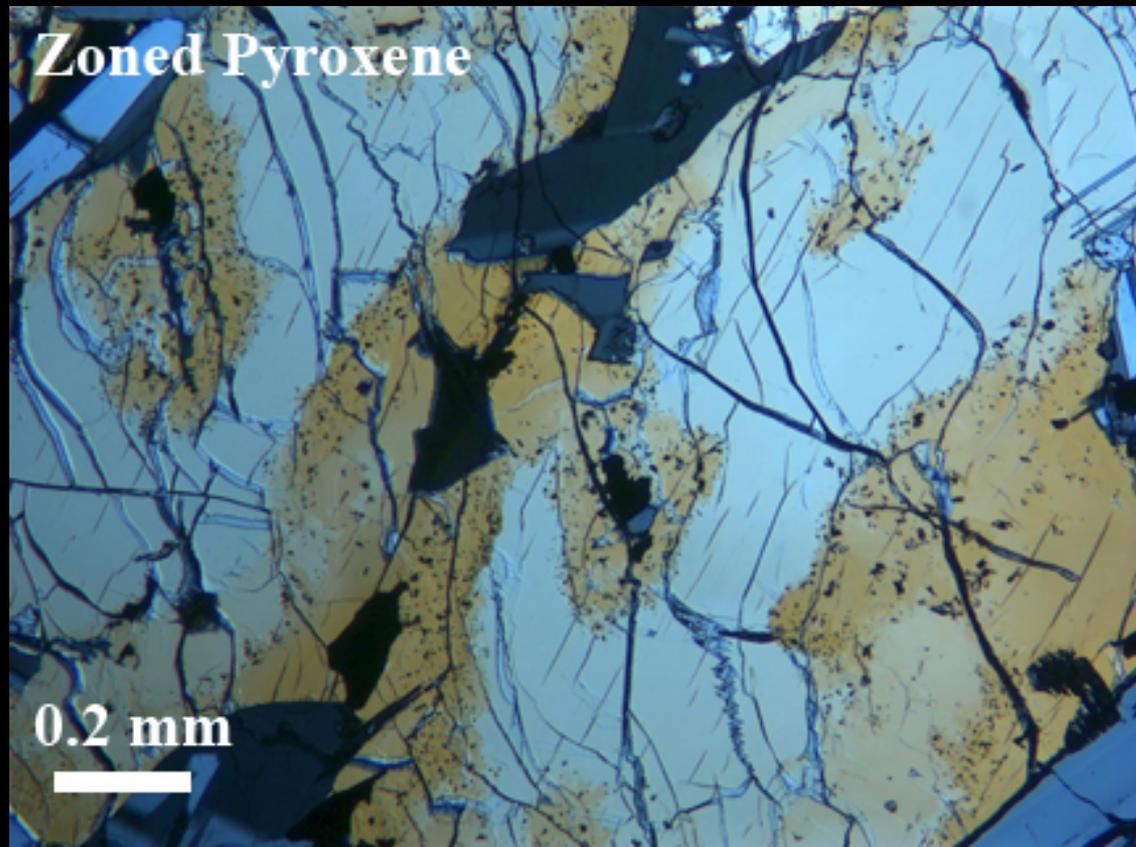
A12 Ilmenite Basalts

- Pyroxene comes on the liquidus early (Papike et al., 1976);
- Stays on the liquidus for much of the crystallization:



On the importance of pyroxene

- It can therefore act as a recording device for most of crystallization;
- Zones within pyroxene provide snapshots of magma at different times.



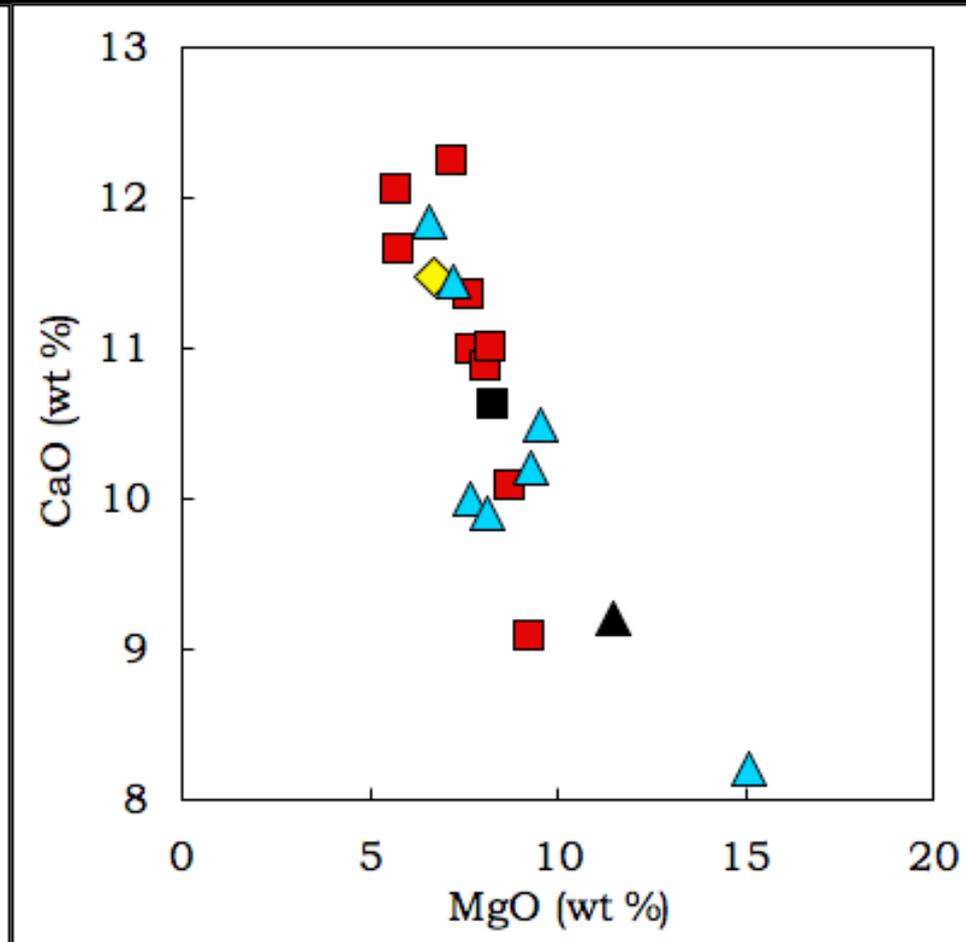
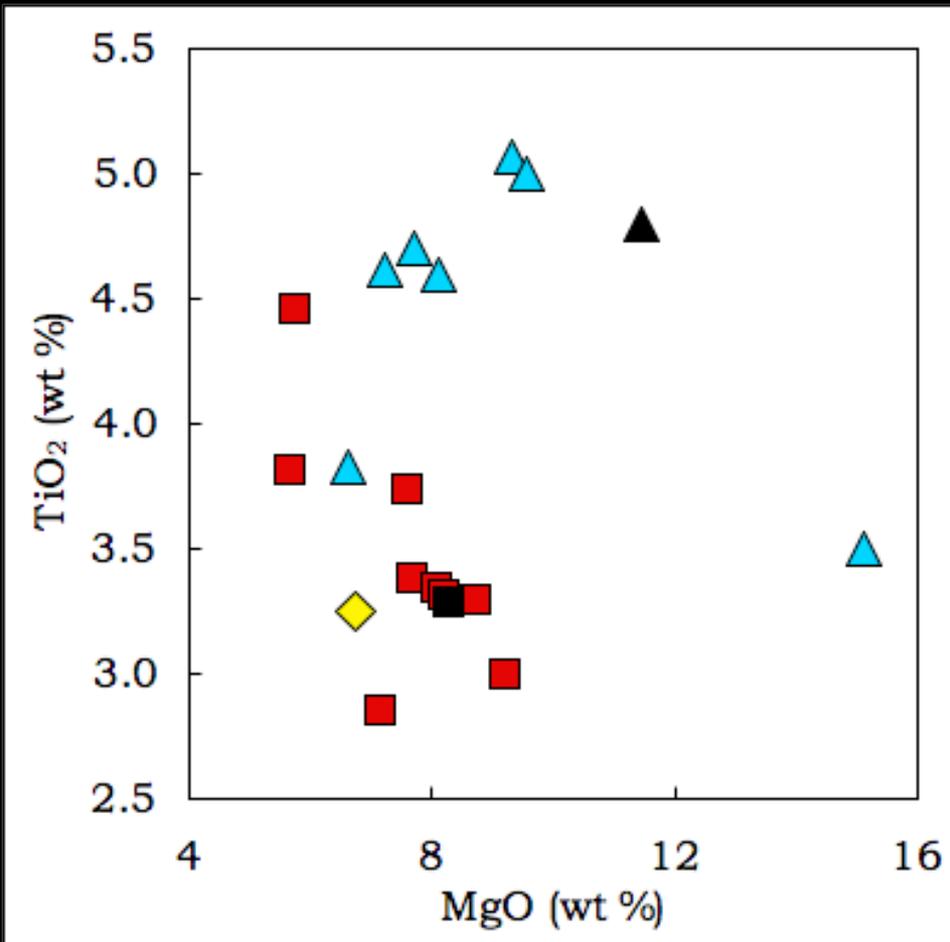
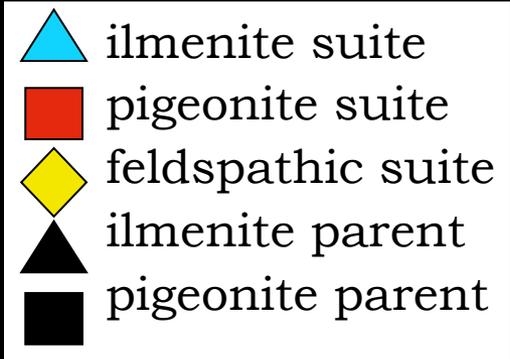
Methods

- Major elemental abundances combined with backscatter images to find primitive pyroxene (*via electron microprobe*)
- We determine trace element concentrations of large primitive pyroxene crystals (*via laser ablation ICPMS*)
- Using published partition coefficients we calculate equilibrium liquids of the basalts

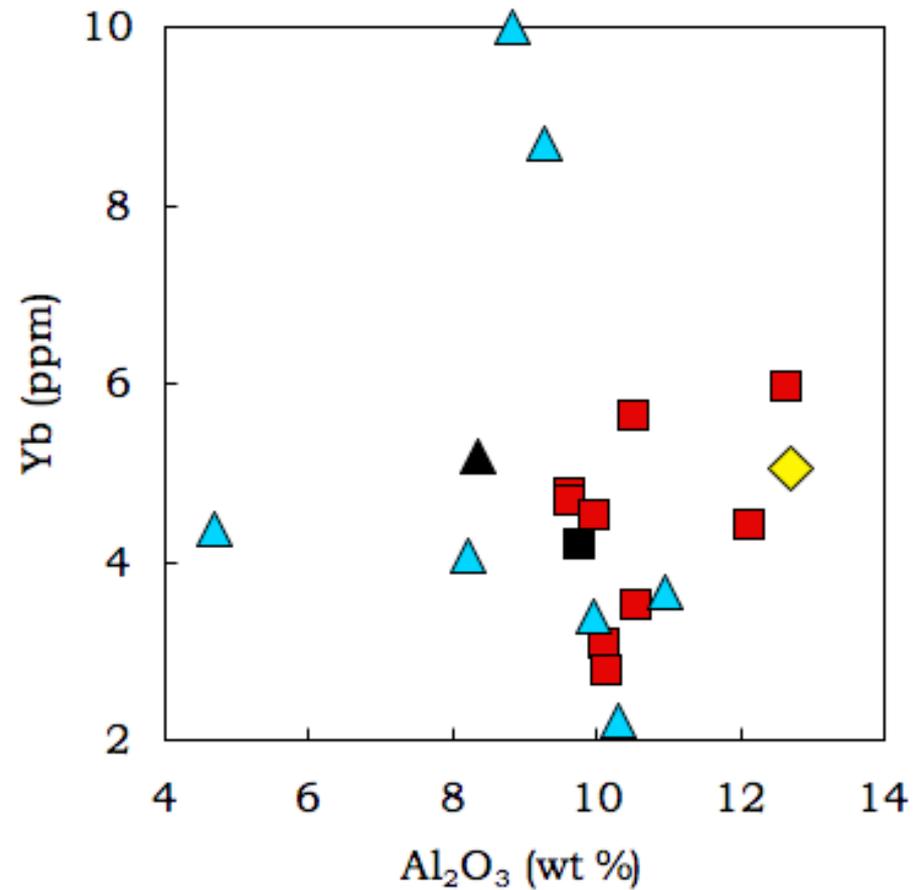
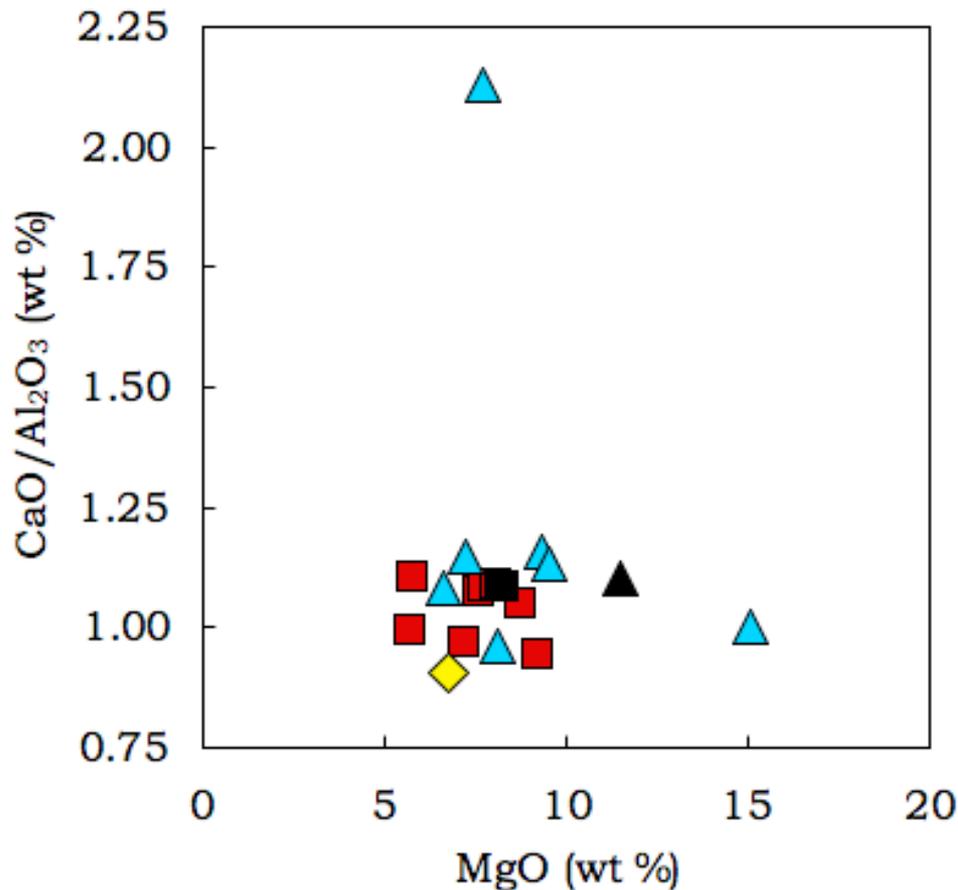
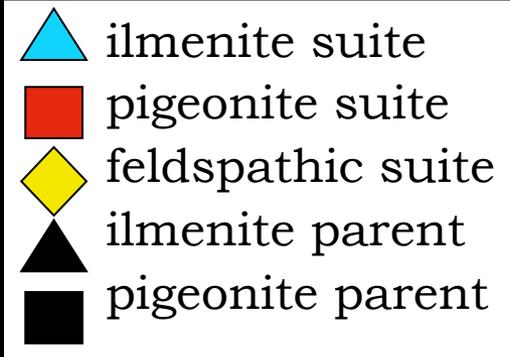
Previously Calculated Apollo 12 Parents

- Neal et al., 1994a,b; used whole rock data from vitrophyric samples to calculate parent for each suite
- Pigeonite Suite- vitrophyre 12011
- Ilmenite Suite- average of vitrophyric/ quenched samples 12008, 12022, 12045
- Feldspathic Suite- not calculated because it is comprised of only one sample

Whole Rock Major Element Trends

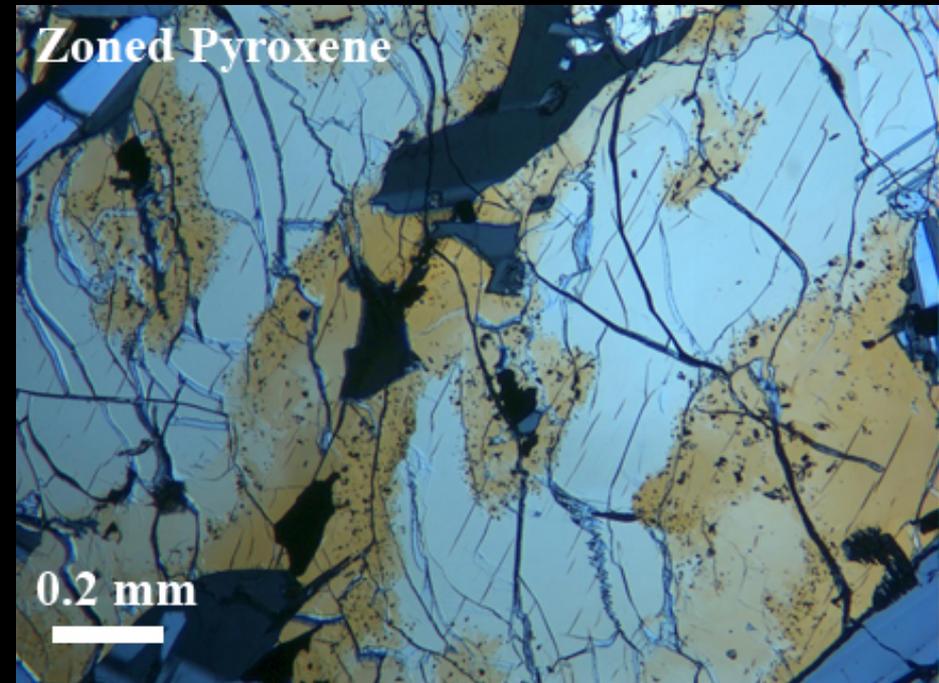


Whole Rock Major Element Trends

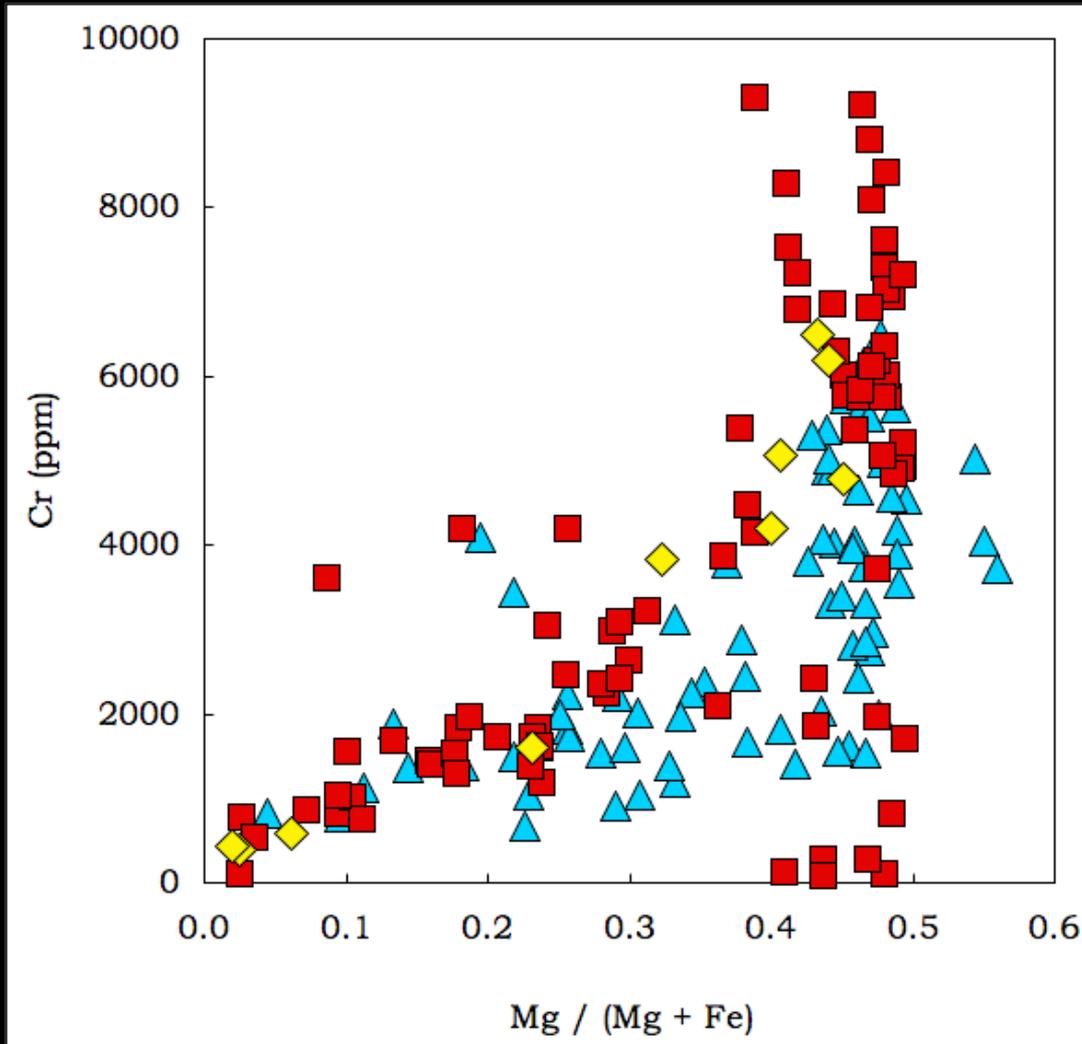


Whole Rock Data

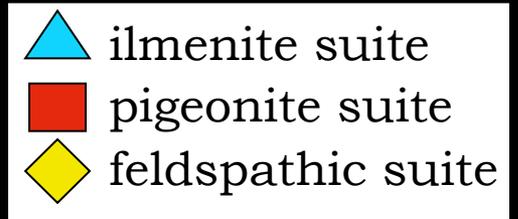
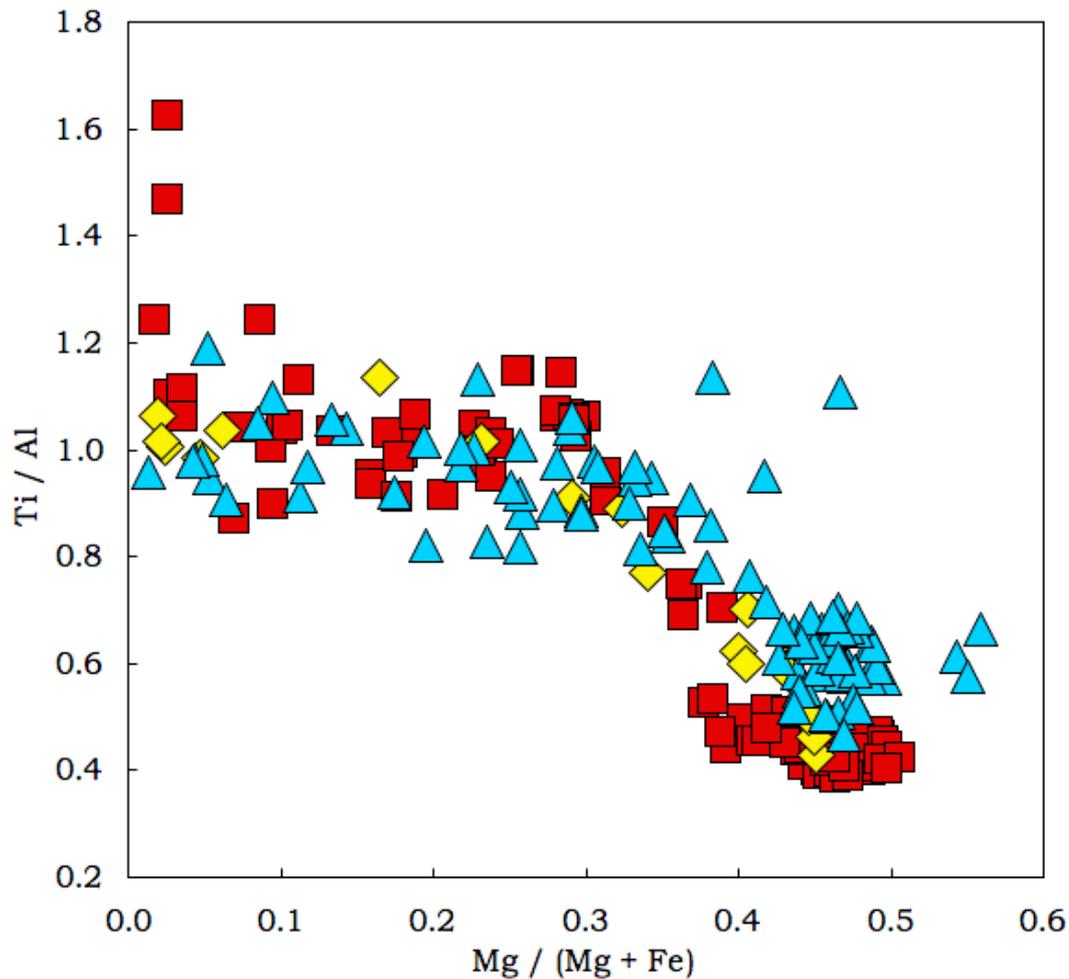
- Whole rock compositions represent an average of the processes that affected a magma during formation/crystallization;
- Crystal stratigraphy (CS) is revealing more primitive compositions than whole rock;
- CS can track magma evolution



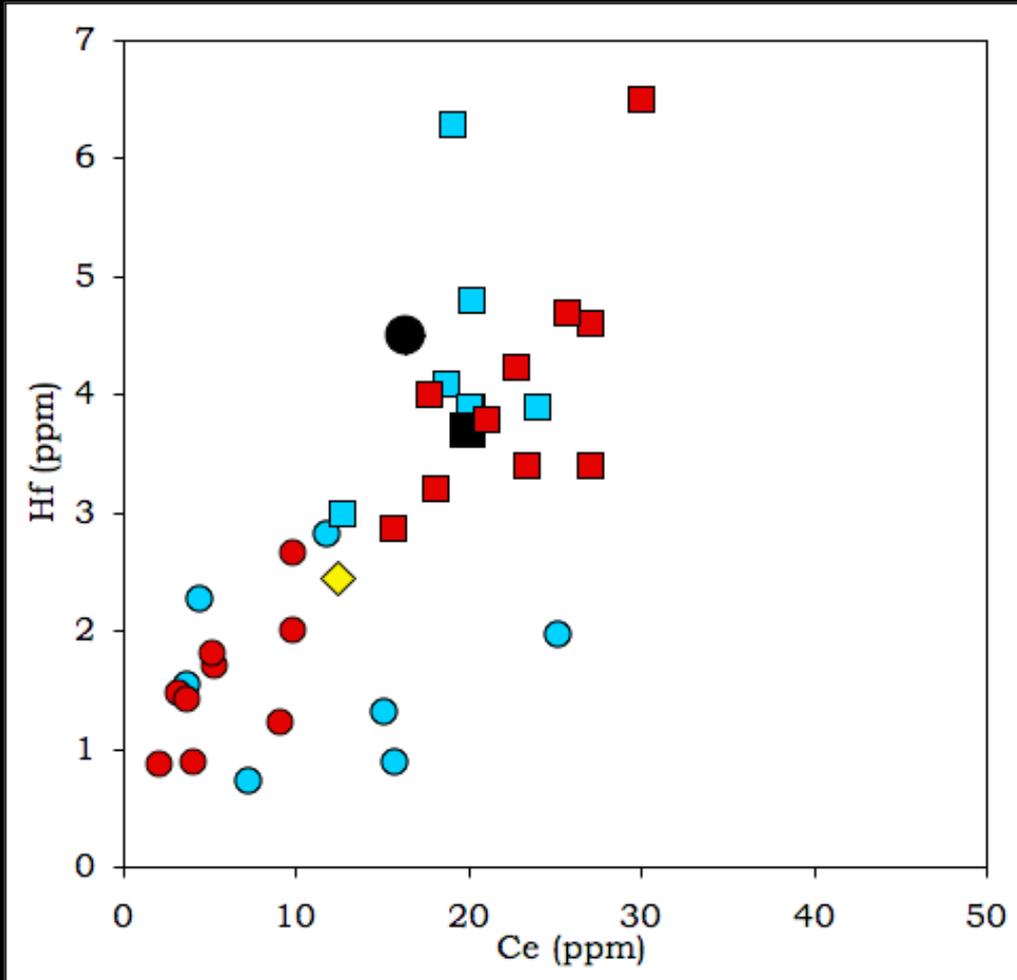
Identifying Primitive Pyroxene



Identifying Primitive Pyroxene

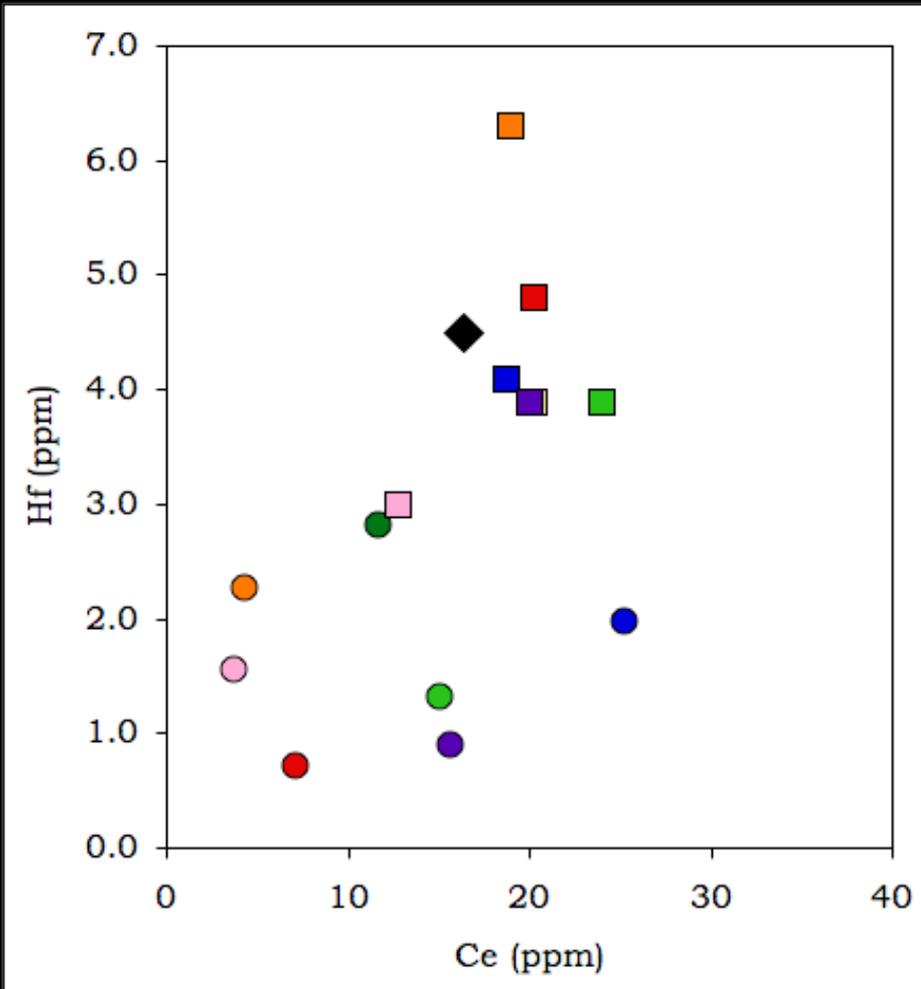


Calculated Equilibrium Liquids



- ilmenite suite- calculated
- ilmenite suite-whole rock
- pigeonite suite- calculated
- pigeonite suite-whole rock
- feldspathic suite- calculated
- ilmenite suite parent
- pigeonite suite parent

Ilmenite Suite Equilibrium Liquids



12016

12062

12051

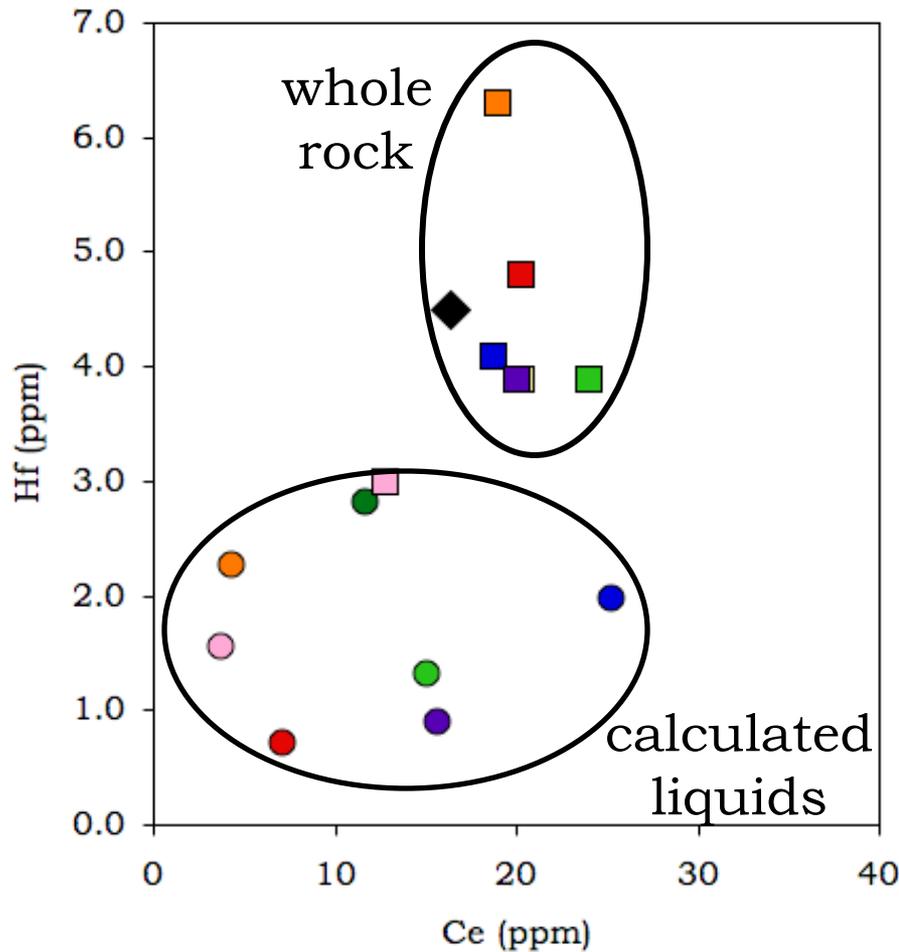
12063

12054

12064

12056

Ilmenite Suite Equilibrium Liquids



12016

12062

12051

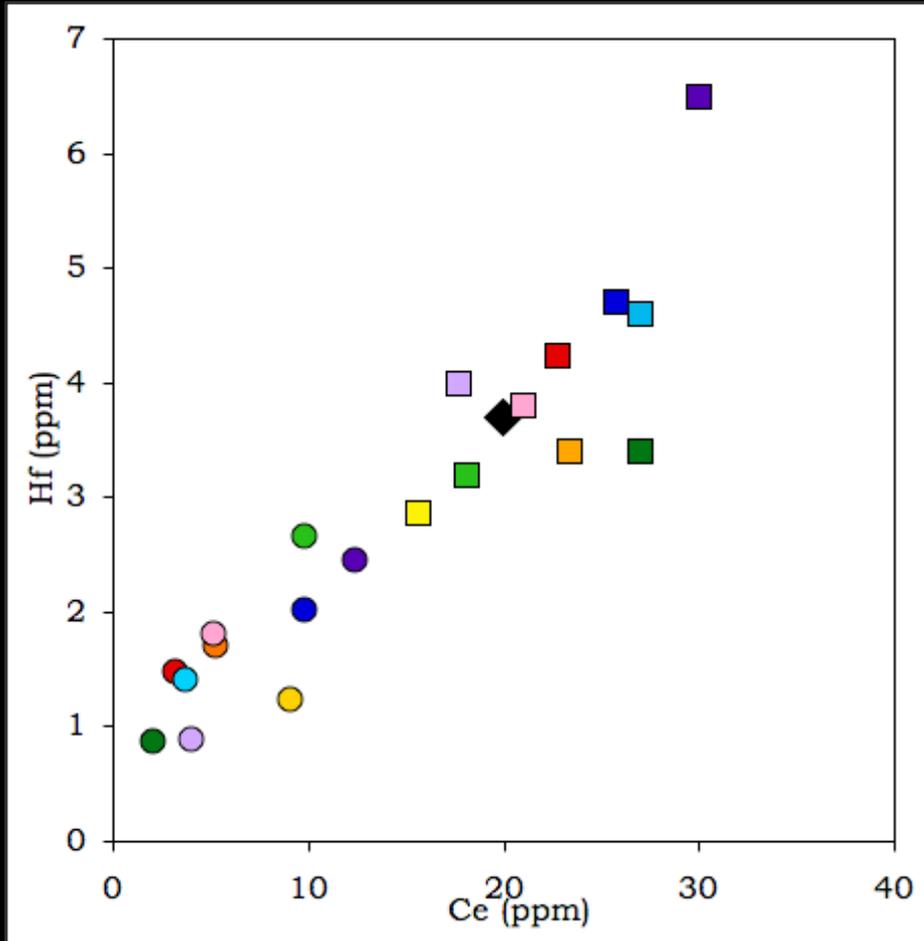
12063

12054

12064

12056

Pigeonite Suite Equilibrium Liquids



12007

12017

12019

12021

12031

12039

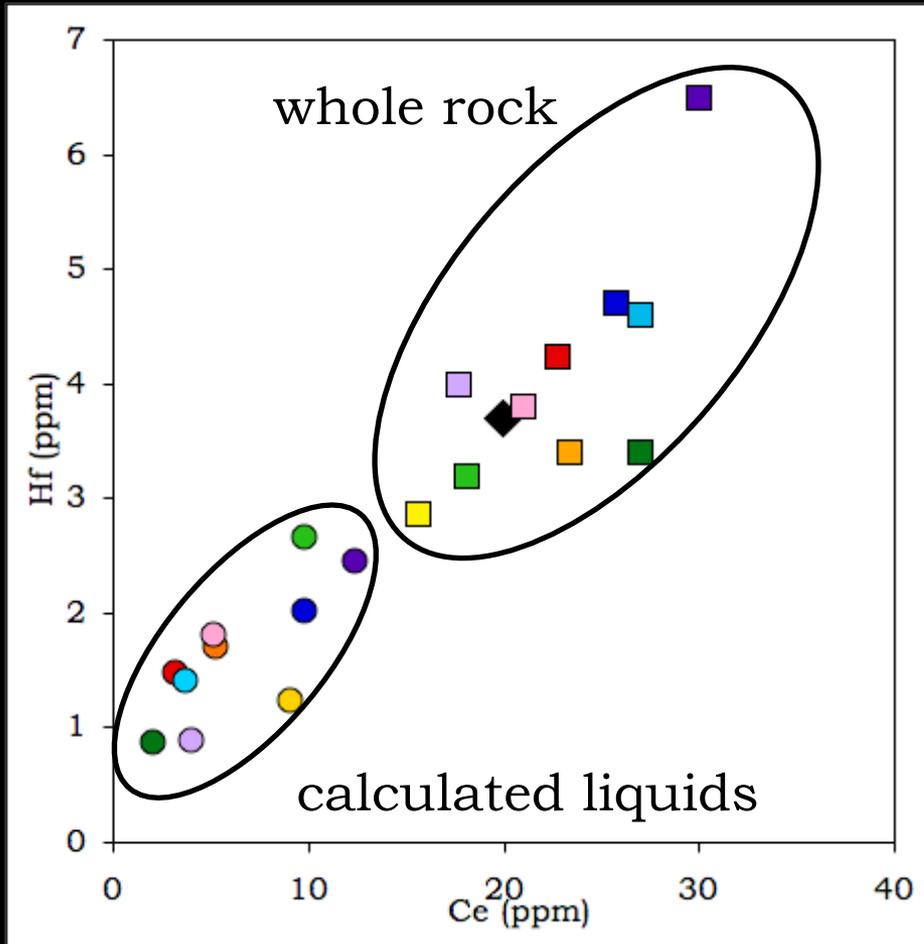
12043

12052

12053

12038

Pigeonite Suite Equilibrium Liquids



12007

12017

12019

12021

12031

12039

12043

12052

12053

12038

Summary & Conclusions

- Crystal stratigraphy is revealing more primitive parental compositions than whole rock data;
- Characterization of parents to limited samples is possible with this method, such as the single Feldspathic suite basalt, 12038;
- The next step is to use these calculated parental liquids to calculate source compositions.

Acknowledgements

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- Thanks to Paul Carpenter and Antonio Simonetti for their guidance in the EMP and ICPMS labs.

