

## Introduction and Motivation

- Previous work suggested that mare basin-related extension on the Moon largely ended ~3.6 Ga<sup>1</sup> and contractional deformation ended ~1.2 Ga<sup>2,3</sup>
- Lunar Reconnaissance Orbiter Camera (LROC) enables the discovery of tectonic landforms at scales not previously imaged<sup>4,5,6</sup>
- Landform morphology<sup>7,8</sup> and stratigraphic relationships imply a complex history of deformation on the Moon

## Landforms

- a. Lobate Scarp: A simple curvilinear, asymmetric hill formed by near-surface fault<sup>4,5,7,8</sup> (Fig. 1a)
- b. Wrinkle Ridge: A complex of curvilinear, asymmetric hills formed by folding over a blind fault or faults<sup>2,9,10</sup> (Fig. 1b)
- c. Graben: A trough formed between two normal faults<sup>1,6</sup> (Fig. 1c)

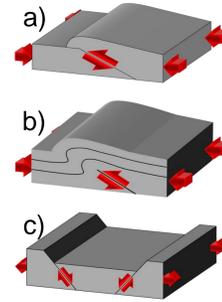


Fig. 1: Block diagrams of a) lobate scarp, b) wrinkle ridge, and c) graben

## Data and Methods

- LROC Narrow Angle Camera (NAC)<sup>11</sup> images with meter-scale resolution
- Nearly continuous NAC image coverage from 45°N to 65°N and 45°W to 45°E
- Map tectonic landforms using ArcGIS both in the mare and in the adjacent highlands
- Identify any preferred orientations for groups of landforms
- Infer principal stress components from surface traces

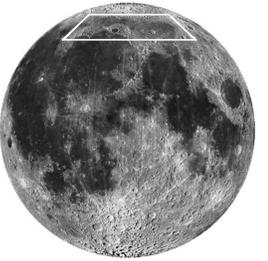


Fig. 2: LROC WAC global orthographic projection centered on 0°N, 0°E showing the location of Mare Frigoris (trapezoid)

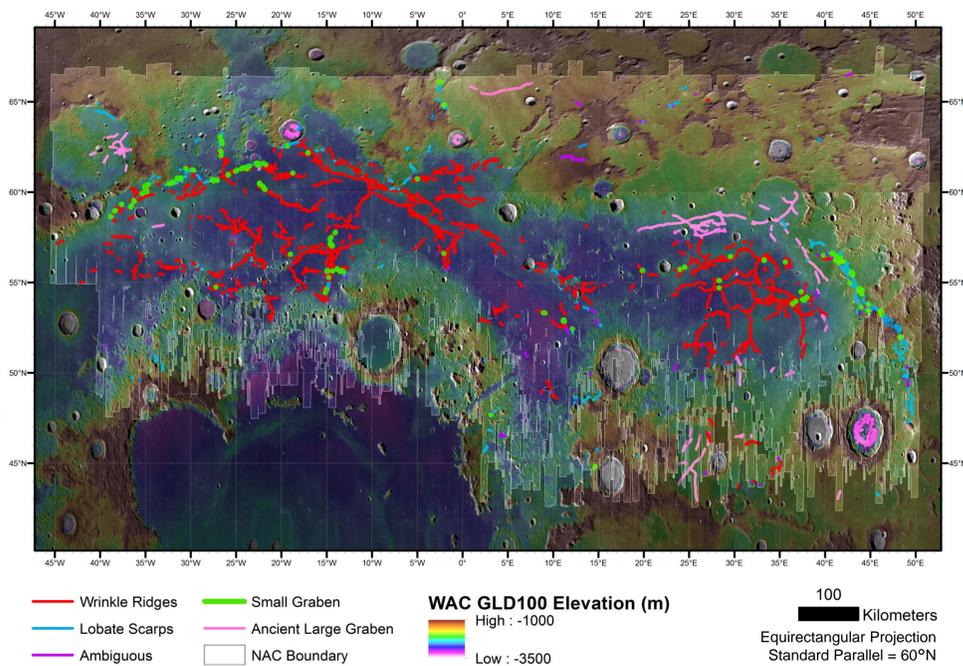


Fig. 3: Tectonic map of Mare Frigoris over LROC WAC DEM<sup>12</sup>

## Perpendicular Graben

- Graben bounded by normal faults should extend in the direction of greatest extensional stress (or least compressive stress,  $\sigma_{min}$ )
- The greatest compressive stress ( $\sigma_{max}$ ) is perpendicular to the direction of slip on a normal fault and parallel to the long axis of the graben
- Conversely for thrust faults,  $\sigma_{max}$  is perpendicular to the strike of the thrust fault and  $\sigma_{min}$  is parallel to its strike
- Small graben perpendicular to ridges/scarps share the same principal stress component orientations and likely formed in the same stress field

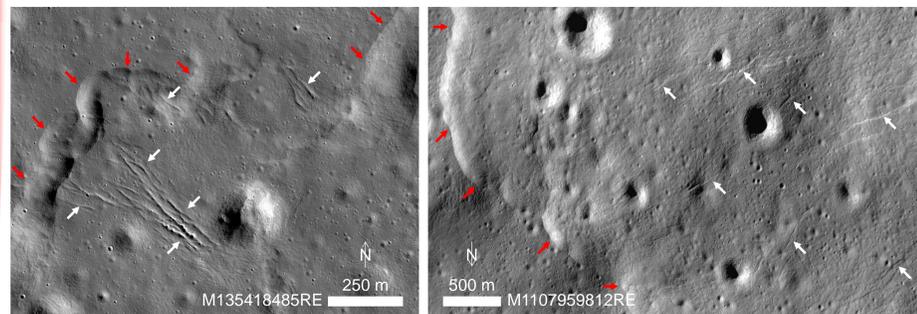


Fig. 4: Small graben (white arrows) nearly perpendicular to a wrinkle ridge (red arrows) at 60.40°N, 34.83°W

Fig. 5: Small graben (white arrows) nearly perpendicular to a lobate scarp (red arrows) at 53.35°N, 46.23°E

## Parallel Graben

- Bending in response to slip on a thrust fault causes an extensional normal stress ( $\sigma_{min}$ ) perpendicular to the fold axis in near surface materials, causing extension parallel to the fault
- Since wrinkle ridges are interpreted as folds overlying thrust faults, we expect some graben to be parallel to ridges, and to a lesser degree graben may also be parallel to lobate scarps as well
- The presence of parallel graben suggests some folding and flexure in Mare Frigoris

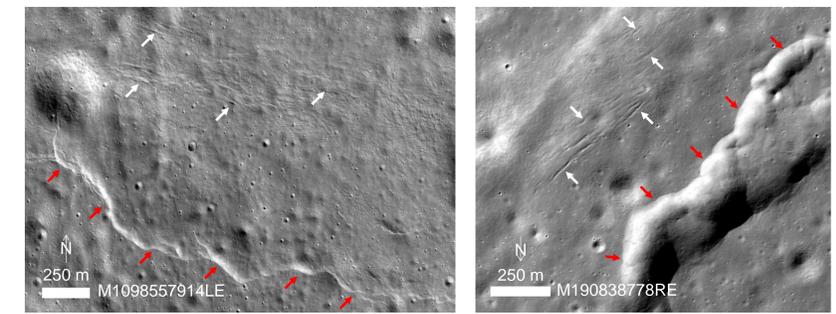


Fig. 6: Small graben (white arrows) nearly parallel to a lobate scarp (red arrows) at 56.14°N, 42.00°E

Fig. 7: Small graben (white arrows) nearly parallel to a wrinkle ridge (red arrows) at 60.03°N, 34.86°W

## Tectonic Map Results

- Over 40 clusters of small graben distributed throughout Mare Frigoris and the neighboring highlands
- Small graben ubiquitously associated with wrinkle ridges and lobate scarps in Mare Frigoris
- Typically only meters to a few tens of meters wide and at most a few meters deep
- Small graben with <1 m of vertical relief should infill in <50 Ma based on analysis of boulder tracks<sup>13</sup>, and if still visible are presumed to be young<sup>6</sup>

## Graben Orientations

- Difference in azimuth measured between small graben and their nearest ridge/scarp
- Graben are preferentially at high angles (nearly perpendicular) to nearby ridges and scarps
- Minor preference for graben to be nearly parallel to ridges and scarps
- Moderate orientations likely due to regolith heterogeneities and curvature of associated ridge/scarp

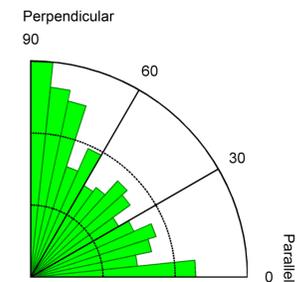


Fig. 8: Length-weighted histogram of azimuthal differences between graben and nearest ridge/scarp

## Conclusions

- Graben nearly perpendicular to nearby wrinkle ridges and lobate share the same principal stress component orientations with those ridges and scarps
- Graben nearly parallel to nearby ridges and scarps suggest folding and flexure in back-limbs of those ridges and scarps
- Small shallow graben are young, perhaps <50 Ma for those with <1 m of vertical displacement and relief
- Very good spatial correlation and consistent stress fields suggest the associated wrinkle ridges and lobate scarps were also likely active within the past 50 Ma
- Tectonic deformation in Mare Frigoris did not completely shut off by ~1.2 Ga, and may continue today

## References

- <sup>1</sup>Lucchitta B.K. and Watkins J.A. (1978) LPS 9, 3459-3472. <sup>2</sup>Hiesinger *et al.* (2003) JGR 108, E001985. <sup>3</sup>Solomon S.C. and Head J.W. (1980) Rev. Geophys. & Space Phys. 18, 107-141. <sup>4</sup>Binder A.B. and Gunga H.C. (1985) Icarus, 63, 421-441. <sup>5</sup>Watters T.R. *et al.* (2010) Science, 329, 936-940. <sup>6</sup>Watters T.R. *et al.* (2012) Nature Geosci., doi:10.1038/ngeo1387. <sup>7</sup>Banks M.E. *et al.* (2012) JGR 117, doi:10.1029/2011JE003907. <sup>8</sup>Williams N.R. *et al.* (2013) JGR 118, 224-233. <sup>9</sup>Schultz R.A. (2000) JGR 105, 12035-12052. <sup>10</sup>Watters T.R. (2004) Icarus 171, 284-294. <sup>11</sup>Robinson M.S., *et al.* (2010) Space Sci. Rev. 150, 81-124. <sup>12</sup>Scholten F. *et al.* (2011) LPSC 42, 2046. <sup>13</sup>Arvidson, R., *et al.* (1975) Moon 13, 67-79.

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